

The importance of accurate a priori information for VLBI Intensive sessions

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1. Introduction
2. Simulation study
3. Impact of erroneous station coordinates
4. Impact of erroneous polar motion
5. Impact of erroneous nutation components
6. Summary & Conclusions

Introduction

Simulation
study

Results UEN

Results PM

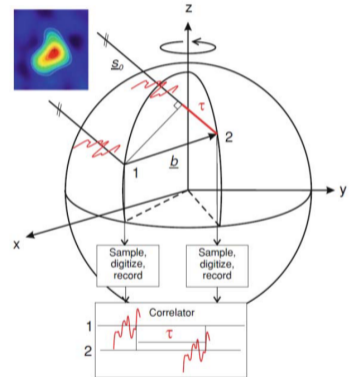
Results NUT

Summary &
Conclusions

- simultaneous measurement of radio signals emitted by extragalactical radio sources (quasars) by at least two VLBI stations

$$\tau = -\frac{\mathbf{b} \cdot \mathbf{s}_0}{c} = t_2 - t_1 \quad (1)$$

- determination of:
 - station and source coordinates,
 - EOP (polar motion, nutation offsets, **UT1-UTC**)
 - atmospheric parameters, ...



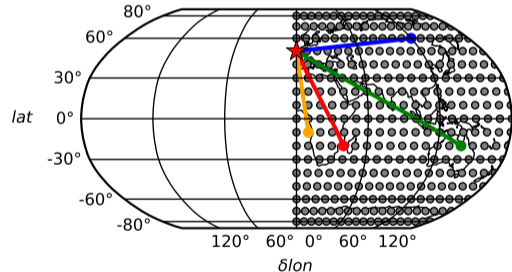
[Schuh and Böhm, 2013]

- single baseline sessions
- one hour duration
- dedicated to derive **UT1-UTC** (mean formal uncertainty: 5-20 μs)
 - restricted number of observations
 - restricted number of estimates (ZWD, clock, UT1-UTC)
 - remaining parameters are **fixed to their a priori value** (EOP, station/ source coordinates, tropospheric gradients, ...)

→ σ_{apriori} **impacts** $\sigma_{\text{UT1-UTC}}$

Experiment setup [Schartner et al., 2021]

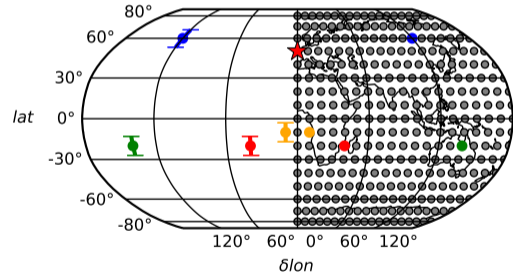
- 323 artificial VGOS stations placed on regular $10^\circ \times 10^\circ$ global grid
- reference stations at $\delta lon = 0 \rightarrow \approx 3000$ baselines
- scheduled (VieSched++) and simulated (VieVS)
 - monthly schedules per baseline
 - reduced source list with equally distributed sources
 - focus corner scheduling algorithm*



*[Nothnagel and Campbell, 1991; Uunila et al., 2012; Gipson and Baver, 2015; Schartner et al., 2021]

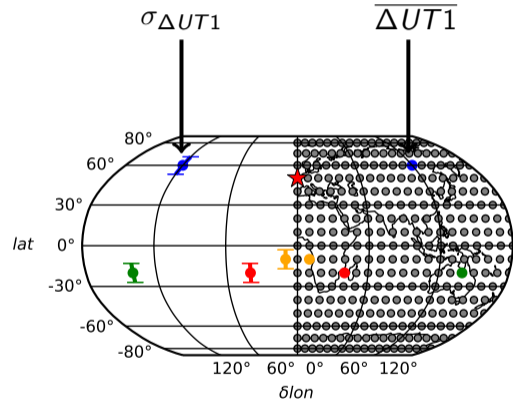
Erroneous a priori information

- introducing realistic errors (in separate evaluations):
 - up, east, north \rightarrow 5 mm
 - $x_p, y_p + dX, dY \rightarrow 162 \mu\text{as}$
- simulation results of evaluations are compared to unaltered results \rightarrow **monthly $\Delta UT1$ values**
- investigate mean and standard deviation of $\Delta UT1$



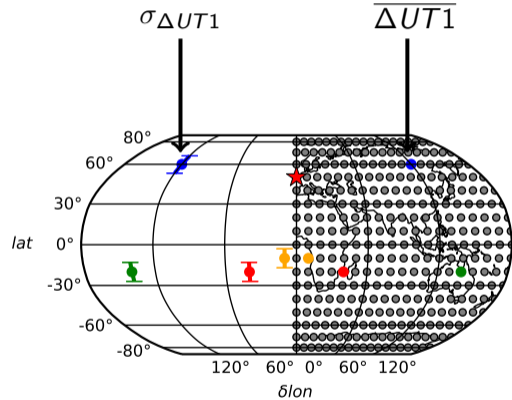
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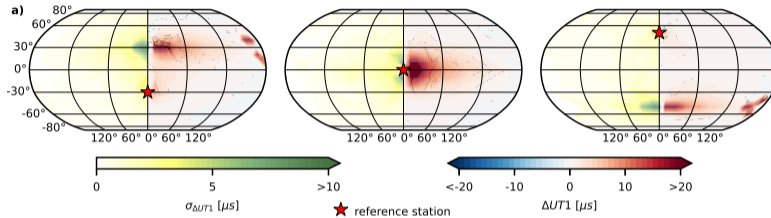


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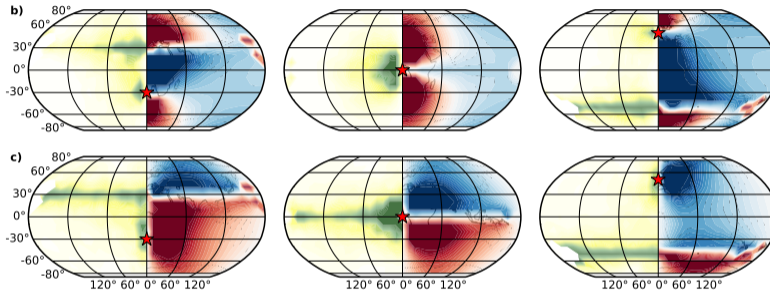


→ high $\overline{\Delta UT1}$... high sensitivity/ low resistance against a priori error
 → high $\sigma_{\Delta UT1}$... high variability of $\overline{\Delta}$ throughout the year



a) error in up-direction

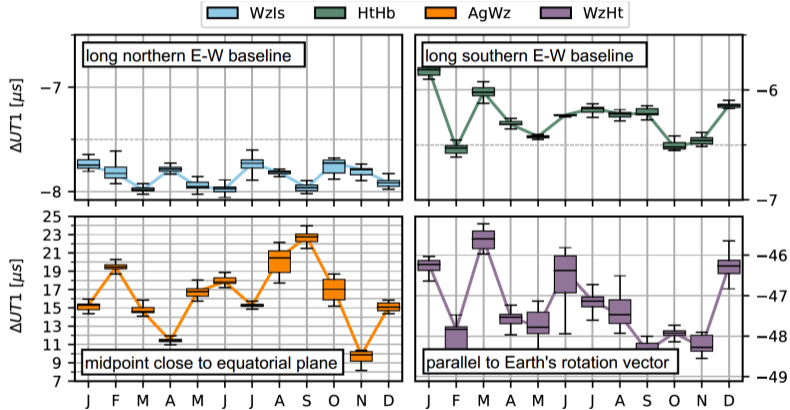
- $\bar{\Delta} > 5 \mu s$ (8%) and $\bar{\Delta} > 20 \mu s$ (3%)
- **low resistance**: baselines with midpoint close to equatorial plane/ baselines parallel to Earth rotation vector



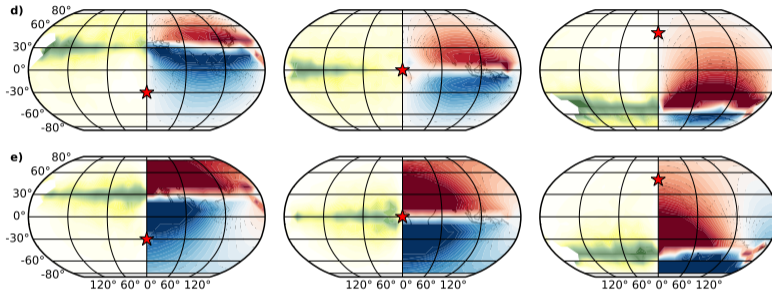
b+c) error in east- and north-direction

- $\bar{\Delta} > 5 \mu\text{s}$ (84/ 63%) and $\bar{\Delta} > 20 \mu\text{s}$ (22/ 16%)
- **low resistance**: baselines with midpoint close to equatorial plane, short baselines (, N-S baselines)
- **high resistance**: E-W baselines to mid-latitudes of same hemisphere

[Kern et al., 2022b - to be published]

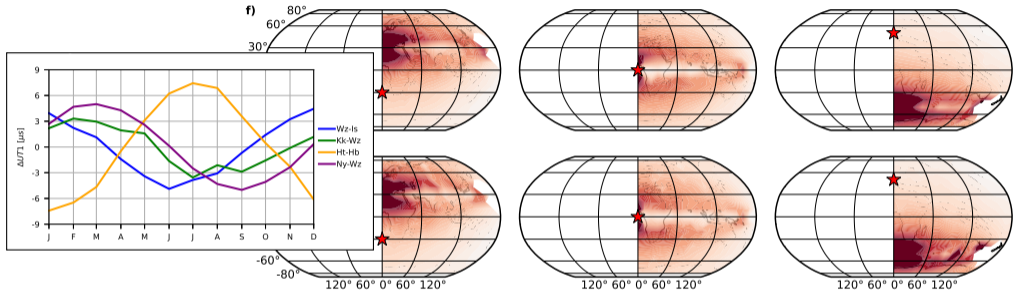


- impact source selection - variability between monthly estimates
- impact scheduling optimization - variations within one month



d+e) error in x_p - and y_p -direction

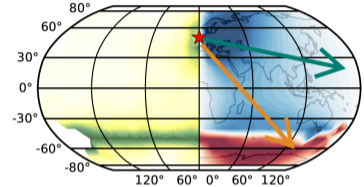
- $\bar{\Delta} > 5 \mu\text{s}$ (50-70%) and $\bar{\Delta} > 20 \mu\text{s}$ (12-25%)
- **low resistance:** baselines with midpoint close to equatorial plane/ equatorial baselines, N-S baselines
- **high resistance:** E-W baselines to mid-latitudes of same hemisphere



f+g) error in dX - and dY -component

- strong dependence on sidereal time \rightarrow amplitude A
- $A > 5 \mu s$ (83%) and $A > 20 \mu s$ (35%)
- **low resistance**: baselines with midpoint close to equatorial plane
- **high resistance**: E-W baselines to mid-latitudes of same hemisphere

- global simulation study on the **impact of a priori errors on the determination of UT1** with VLBI Intensives
- almost 3000 baselines and 240 000 simulations



high sensitivity/ low resistance against investigated a priori errors

baselines with a midpoint close to equatorial plane

low sensitivity/ high resistance against investigated a priori errors

long E-W baselines between a reference station and a station at mid-latitudes of same hemisphere

- impact of a priori errors are **not negligible!**

Gipson J, Baver K (2015) Minimization of the UT1 Formal Error Through a Minimization Algorithm. In: Proceedings of the 22nd European VLBI Group for Geodesy and Astrometry Working Meeting

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Nothnagel A, Campbell J (1991) Polar Motion Observed by Daily VLBI Measurements; Proceedings of the AGU Chapman Conference on Geodetic VLBI: Monitoring Global Change; NOAA Technical Report NOS 137 NGS 49, S. 345 - 354, Washington D.C.

Schartner M, Kern L, Nothnagel A, Böhm J, Soja B (2021) Optimal VLBI baseline geometry for UT1-UTC Intensive observations. Journal of Geodesy 95(75)

Schuh H, Böhm J (2013) Very Long Baseline Interferometry for Geodesy and Astrometry. In G. Xu, editor, Sciences of Geodesy - II, pages 339–376. Springer-Verlag Berlin Heidelberg

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